# Towards Unified Approaches to Nuclear Structure and Reactions for Nuclei

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## **Collaborators**

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C. Robin (Uni. of Washington)

Exact

Ab initio

M. Dupuis (CEA)

V. Zelevinsky (MSU)

Configuration

interaction



density functional



Lawrence Livermore National Laboratory



### **MOTIVATIONS: DESCRIPTION OF EXOTIC (NEUTRON RICH) AND RESONANT SCATTERING**

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### TRADITIONAL BOUND STATE TECHNIQUES

Methods developped in this presentation to solve the many body problem:





### EQUAL TREATMENT OF BOUND AND RESONANT STATES IN THE AB INITIO NO-CORE SHELL MODEL WITH CONTINUUM

Methods developped in this presentation to solve the many body problem:



#### **UNDERSTANDING NUCLEAR STRUCTURE: 6Li**





### UNDERSTANDING NUCLEAR STRUCTURE: <sup>6</sup>Li







### UNDERSTANDING NUCLEAR STRUCTURE: 6Li





- Consistent *ab initio* calculation of this system.
- Bound and resonant states are treated in the same way.
- Effects of the chiral **3N force** are revealed.



# CONCURRENTLY WITH NUCLEAR REACTION $^{2}H(\alpha, d)^{4}He$





# p-1°C SCATTERING: STRUCTURE OF 11N RESONANCES & TESTING $\chi$ HAMILTONIAN





A. Kumar, R. Kanungo, A. Sanetullaev *et al.* 



### *n*-<sup>4</sup>He SCATTERING: NN VERSUS 3N INTERACTIONS

G. Hupin, J. Langhammer et al., PRC88 (2013); P. Navrátil, S. Quaglioni, G. Hupin et al., Phys. Scri.91 (2016) Celebrating the 1975 Nobel Prize





- The 3N interactions influence mostly the *P* waves.
- The largest splitting between *P* waves is obtained with NN+3N.

Comparison between NN+3Nind and NN+3N at N<sub>max</sub>=13 with six <sup>4</sup>He states and 14 <sup>5</sup>He states.



# *n*-<sup>4</sup>He SCATTERING: *AB INITIO* VS EXACT





• **Good agreement** between the two methods.



# LIGHT SYSTEMS, FIRST TASTE OF EXOTIC PHENOMENA

ХН
 ХН
 АК
 ЖНХ-ТЧИА
 ЖНХ-ТЧИА

- Bound states
- Resonant states
- Scattering states
- Halo densities
- Clusterizations

•

...

One-Proton Halo			Stable			<sup>17</sup> Ne	<sup>18</sup> Ne	<sup>19</sup> Ne	<sup>20</sup> Ne				
Two-Proton Halo						<sup>15</sup> F	<sup>16</sup> F	<sup>17</sup> F	18F	<sup>19</sup> F	•	<b>3</b>	
One-Neutron Halo				<sup>12</sup> O	130	<sup>14</sup> O	150	<sup>16</sup> O	<sup>17</sup> O	<sup>18</sup> O			
Two-Neutron Halo				<sup>11</sup> N	<sup>12</sup> N	<sup>13</sup> N	<sup>14</sup> N	<sup>15</sup> N	<sup>16</sup> N	<sup>17</sup> N	<sup>18</sup> N	<sup>19</sup> N	<sup>20</sup> N
Unbound <sup>9</sup> C			10C	11C	<sup>12</sup> C	<sup>13</sup> C	<sup>14</sup> C	15C	<sup>16</sup> C	17C	<sup>18</sup> C	<sup>19</sup> C	
2 ∎ <sup>8</sup> B			9B	<sup>10</sup> B	<sup>11</sup> B	<sup>12</sup> B	<sup>13</sup> B	<sup>14</sup> B	<sup>15</sup> B	<sup>16</sup> B	<sup>17</sup> B	<sup>18</sup> B	
<sup>7</sup> Be			<sup>7</sup> Be	<sup>8</sup> Be	<sup>9</sup> Be	<sup>10</sup> Be	<sup>11</sup> Be	<sup>12</sup> Be	<sup>13</sup> Be	<sup>14</sup> Be	<sup>15</sup> Be	<sup>16</sup> Be	<sup>17</sup> Be
		<sup>5</sup> Li	<sup>6</sup> Li	<sup>7</sup> Li	<sup>8</sup> Li	<sup>9</sup> Li	<sup>10</sup> Li	<sup>11</sup> Li	<sup>12</sup> Li	<sup>13</sup> Li			
	<sup>3</sup> He	⁴He	<sup>5</sup> He	<sup>6</sup> He	<sup>7</sup> He	<sup>8</sup> He	<sup>9</sup> He	<sup>10</sup> He					
11	H <sup>2</sup> H	ЗН	⁴H	⁵H	6H	7H			ľ	<b>8</b>			
	n		→ N ●										



# LIGHT SYSTEMS, FIRST TASTE OF EXOTIC PHENOMENA

8 SRG-N<sup>3</sup>LO NN <sup>6</sup>He Stable One-Proton Halo  $E_{th}(\alpha+n+n)$  (MeV) 66 **Two-Proton Halo** +--**One-Neutron Halo** 120 130 ХЫЧМ Ľ. Two-Neutron Halo 11N 12N 이신  $\alpha + n + n$ Unbound 9C 10C 11C NCSMC NCSMC Expt. NCSM (∞)  $\Lambda$ =1.5 fm<sup>-1</sup> A=2.0 fm<sup>-1</sup> <sup>8</sup>B 9B 10B 2.6 2.6 Viewek(N<sup>3</sup>LO)  $SRG-N^3LONN, \Lambda = 2.0 \text{ fm}^3$ 2.4 2.4rms Radius (fm) ...... <sup>8</sup>Be 9Be 7Be 2.22.2 <sup>6</sup>He **Bound states**  $\bullet$ 2.0 2.08Li 5Li 6Li **Resonant states** 1.8 1.8  $\bullet$ 4He <sup>3</sup>He 5 6He е **Scattering states**  $\bullet$ 15 **Halo densities** • 1.0 $^{1}H$  $^{2}H$ <sup>3</sup>H 1 1 × 100 4H 6H 0.5 **Clusterizations**  $\bullet$ Exp. Th. NCSMC 0.0 0.0 n r\_\_\_ NCSM • ... r<sub>pp</sub> -0.5 ··· EIHH S20 mm -1.0-11-00 4 6 8 10 1.8 2.0 $\Lambda_{kwk}(fm^{-1})$  $N_{\rm max}$ 



### LOW-ENERGY TRANSFER REACTIONS (d, N)

Primordial Nucleosynthesis (blue)  $(\alpha, \gamma)$ (0.,n)  $(\mathbf{c})$ \*C (β-) Έ (p,)) (t, j) 77 10B) (12B) (11B)х  $(n, \gamma)$ Be (B+) °Li Li 'Li (d,γ) # (t,n) (He) (He) (d,n) 🔿 (t,p) х 2 H (p,α) H (d,p) ké (n,p)  $(d,n\alpha)$ n (n,α)



### LOW-ENERGY TRANSFER REACTIONS (d, N)



La science doit « nous rendre comme maîtres et possesseurs de la Nature » R. Descartes *Discours de la méthode*.

#### ITER design (Cadarache, France)





### **DT POLARIZED THERMONUCLEAR FUSION**





### **DT POLARIZED THERMONUCLEAR FUSION**





- Importance of structure of neighboring resonances is revealed in transfer reactions.
- **Predictions** for  ${}^{3}\vec{\mathrm{H}}(\vec{d},n){}^{4}\mathrm{He}$  reaction and its enhancement factor.



# **POLARIZED THERMONUCLEAR FUSION:**

Angular distribution in different polarization scenarios J = 1/2I = 1 $J_z = -1/2$  $J_z=1$ Total cross section increased (on average) No changes Total cross section decreased p\_=0.8, p\_.=0.8 0.47500.4323Unpolarized - -0.675 0.4318 0.3783 $\begin{bmatrix} 0.3183\\ -1.187\\ 0.3242\\ -0.2702\\ 0.2702\\ 0 \end{bmatrix}$  $\left[ 1 - 133 + 0.3886 \right]$  $\frac{\partial \sigma}{\partial \theta} \left[ \mathrm{mb.sr}^{-1} \right]$ 0.540 $p_z=0.8, p_z^t=-0.8$  $p_{zz} = 0.8$ 0.405Unpolarized 0.270 $p_{:}=0.8, p_{:}^{t}=0.8$ 0.3023  $p_{11}=0.8$ 0.2161 Unpolarized 0.135120 30 90 1503060 901201503060 901201500 60 0.  $\theta_{\rm c.m.}$  [deg]  $\theta_{\rm c.m.}$  [deg]  $\theta_{\rm c.m.}$  [deg]

Spin tensor properties of the deuteron give the angular shape. (Same as in  ${}^{3}\overrightarrow{He}(\overrightarrow{d},p){}^{4}He$ )



N. Pillet, J.-F. Berger, and E. Caurier, PRC78 (2008); C. Robin, N. Pillet, D. Peña Arteaga, and J.-F. Berger, PRC83 (2016)





$$\Psi_{MPMH}^{(A)} = \sum_{\alpha} c_{\alpha} \mathcal{A}_{\alpha} \varphi_{\alpha,A}(\vec{r}_{A}) \varphi_{\alpha,A-1}(\vec{r}_{A-1}) \dots \varphi_{\alpha,1}(\vec{r}_{1})$$

$$\underset{\text{Mixing coefficients}(unknown)}{\text{Single particle states}(unknown)}$$

# For computational purpose we can truncate further

✓ Based on many-body energy:

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Based on an active valence space and inactive core:





✓ According to the type of many-body excitations



- All nuclear correlations thought to be important in mean-field can be included.
- Some key features :
  - Systematically improvable, i.e. all the Hilbert space can be spanned.
  - All symmetries but translational invariance are conserved.





### MPMH METHOD: ONE OR TWO PROPERTIES OF INTEREST

C. Robin, N. Pillet, D. Peña Arteaga, and J.-F. Berger PRC83 (2016)

Variation of mixing coefficients: large-scale diagonalization.

Variation of orbitals: generalized **one-body problem** embedded **in manybody space** spanned by Ψ.

Mean-field adapted to a given type of many-body correlations (NCSMC, Shell-Model, 2p-2h...).

$$\delta_{c_{\alpha}^{*}} (E[\Psi] - \lambda(\langle \Psi | \Psi \rangle - 1)) = 0$$
$$\Leftrightarrow \sum_{\beta} \langle \Phi_{\alpha} | H | \Phi_{\beta} \rangle c_{\beta} = \lambda c_{\alpha}$$

$$\delta_{\varphi_i^*} (E[\Psi] - \lambda(\langle \Psi | \Psi \rangle - 1)) = 0$$
  
$$\Leftrightarrow [h[\rho], \rho] = G[\sigma]$$

Orbitals are optimum:  $|\varphi\rangle = U|\varphi\rangle$ 





### **CONNECTING EDF TO BARE INTERACTION WITH MPMH**

### Ab initio $\rightarrow$ Mean field

#### use MPMH for the first time:

 ➢ To renormalize a bare nuclear interaction and reveal the corresponding mean-field.
 → MPMH can treat beyond mean-field correlations.







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- $\rightarrow$  Need for a more general functional form (yet easy to integrate).

$$v_{12}[\rho] = \sum_{j=1}^{2} (W_j + B_j P_{\sigma} - H_j P_{\tau} - M_j P_{\sigma} P_{\tau}) e^{-(\vec{r}_1 - \vec{r}_2)^2 / \mu_j} + (W_j + B_j P_{\sigma} - H_j P_{\tau} - M_j P_{\sigma} P_{\tau}) e^{-(\vec{r}_1 - \vec{r}_2)^2 / \mu_3} \frac{\rho^{\alpha}(r_1) + \rho^{\alpha}(r_2)}{2(\mu_3 \pi)^{3/2}} + i W_{LS} \overline{\nabla}_{12} \delta(\vec{r}_1 - \vec{r}_2) \wedge \overline{\nabla}_{12} (\vec{\sigma}_1 + \vec{\sigma}_2) + \text{tensor}$$







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+  $(W_j + B_j P_{\sigma} - H_j P_{\tau} - M_j P_{\sigma} P_{\tau}) e^{-(\vec{r}_1 - \vec{r}_2)^2 / \mu_3} \frac{\rho^{\alpha}(r_1) + \rho^{\alpha}(r_2)}{2(\mu_3 \pi)^{3/2}}$   
+  $iW_{LS} \overline{\nabla}_{12} \delta(\vec{r}_1 - \vec{r}_2) \wedge \overline{\nabla}_{12} (\vec{\sigma}_1 + \vec{\sigma}_2)$   
+ tensor





Use properties of the continuum to infer an interaction fitted for exotic systems: a first step towards reactions.

Continuum properties of <sup>11</sup> N											
	NN-	-3N400	$N^{2}L$	$O_{\rm sat}$	Data evaluation [12]						
$J^{\pi}$	$E_r$	Г	$E_r$	Г	$E_r$	Γ					
$1/2^{+}$	1.47	5.53	1.33	1.45	1.49(6)	0.83(3)					
$1/2^{-}$	1.91	0.55	1.95	0.57	2.22(3)	0.6(1)					



### **IN DEVELOPMENT: RESONANCES**



**Aguilar-Balslev-Combes theorem**: the resonant states of the original Hamiltonian are invariant and the non-resonant scattering states are rotated and distributed on a  $2\theta$  ray that cuts the complex energy plane with a corresponding threshold being the rotation point.

$$\widehat{H}(r,\theta)\psi(r,\theta) = (E+i\Gamma)\psi(r,\theta)$$
  
Energy 
$$\begin{array}{c} & & \\$$



### **IN DEVELOPMENT: RESONANCES**









# **IN DEVELOPMENT: APPLICATION TO NN SYSTEM**

 ${}^{1}S_{0} \,\mathrm{N}^{3}\,\mathrm{LO}\,\mathrm{pn}\,\mathrm{interaction}\,\mathrm{at}\,\theta = 10.0^{\circ}$ 



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- SRG evolution requires spanning a large NN basis  $(n_r \sim 150)$ . The typical scale of k is  $10 \text{ fm}^{-1}$   $(V^{NN})$ .
- Complex scaling involves the integration of diverging polynomials (of order n) far from their zeroes.



# SOFT INTERACTION FOR CI METHODS WITH SRG

 ${}^{1}S_{0}$  N<sup>3</sup>LO pn interaction at  $\theta = 10.0^{\circ} \Lambda_{\text{SRG}} = 1.5 \text{ fm}^{-1}$ 







#### **THANK YOU!**





2 4 6 8 10 neutron's separation (fm)

<sup>4</sup>He-neutrons separation (fm)

0

2

<sup>6</sup>He

Cigar